## The TT metric GW worksheet

Edoardo Milotti

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To carry out the calculations we need to recall a few important expressions:

• Christoffel symbols

$$\Gamma^{\mu}_{\alpha\beta} = \frac{1}{2} g^{\mu\nu} \left( \partial_{\alpha} g_{\nu\beta} + \partial_{\beta} g_{\nu\alpha} - \partial_{\nu} g_{\alpha\beta} \right) \tag{1}$$

 $\bullet\,$  Riemann tensor

$$R^{\mu}_{\alpha\beta\gamma} = \partial_{\beta}\Gamma^{\mu}_{\alpha\gamma} - \partial_{\gamma}\Gamma^{\mu}_{\alpha\beta} + \Gamma^{\delta}_{\alpha\gamma}\Gamma^{\mu}_{\delta\beta} - \Gamma^{\delta}_{\alpha\beta}\Gamma^{\mu}_{\delta\gamma}$$
(2)

• Ricci tensor

$$R_{\alpha\beta} = R^{\mu}_{\alpha\beta\mu} = \partial_{\beta}\Gamma^{\mu}_{\alpha\mu} - \partial_{\mu}\Gamma^{\mu}_{\alpha\beta} + \Gamma^{\delta}_{\alpha\mu}\Gamma^{\mu}_{\delta\beta} - \Gamma^{\delta}_{\alpha\beta}\Gamma^{\mu}_{\delta\mu}$$
(3)

 $\bullet\,$  metric tensor for the + polarization

$$[g_{\mu\nu}] = [\eta^{\mu\nu}] + [h^{\mu\nu}] = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & -1 + h(t,z) & 0 & 0 \\ 0 & 0 & -1 - h(t,z) & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$
(4)

• inverse metric tensor for the + polarization

$$[g^{\mu\nu}] = [\eta^{\mu\nu}] - [h^{\mu\nu}] = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & -1 - h(t,z) & 0 & 0 \\ 0 & 0 & -1 + h(t,z) & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}$$
(5)

Next, assuming a sinusoidal time dependence

$$h(t,z) = A\cos(\omega t - kz) = h_{xx}^{TT} = -h_{yy}^{TT}$$
 (6)

and using the *Diagonal metric worksheet* of T. A. Moore<sup>1</sup> we carry out the calculations that show that for the + polarization

$$R_{tt} = -\frac{1}{c^2} \left( h\ddot{h} + \frac{\dot{h}^2}{2} \right); \quad R_{zz} = R_{tt}; \quad R_{xx} = R_{yy} = 0$$
(7)

to second order in h.

After averaging over time, we find

$$-\left\langle h\ddot{h} + \frac{\dot{h}^2}{2} \right\rangle = \omega^2 A^2 \left\langle \cos^2(\omega t - kz) - \frac{1}{2}\sin^2(\omega t - kz) \right\rangle$$
$$= \omega^2 A^2 \left\langle \cos(2\omega t - 2kz) \right\rangle + \frac{1}{2} \left\langle \dot{h}^2 \right\rangle = \frac{1}{2} \left\langle \dot{h}^2 \right\rangle \quad (8)$$

Finally, it is easy to see that the *h*-dependent part of the Ricci scalar vanishes.

<sup>&</sup>lt;sup>1</sup>We must be careful that in Moore's text he uses the (-,+,+,+) signature while we use the (+,-,-,-) signature, and units such that c = 1. He also uses a different convention for the Ricci tensor, so that the sign is reversed with respect to our definition.