

# The TT metric GW worksheet

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To carry out the calculations we need to recall a few important expressions:

- Christoffel symbols

$$\Gamma_{\alpha\beta}^{\mu} = \frac{1}{2}g^{\mu\nu} (\partial_{\alpha}g_{\nu\beta} + \partial_{\beta}g_{\nu\alpha} - \partial_{\nu}g_{\alpha\beta}) \quad (1)$$

- Riemann tensor

$$R_{\alpha\beta\gamma}^{\mu} = \partial_{\beta}\Gamma_{\alpha\gamma}^{\mu} - \partial_{\gamma}\Gamma_{\alpha\beta}^{\mu} + \Gamma_{\alpha\gamma}^{\delta}\Gamma_{\delta\beta}^{\mu} - \Gamma_{\alpha\beta}^{\delta}\Gamma_{\delta\gamma}^{\mu} \quad (2)$$

- Ricci tensor

$$R_{\alpha\beta} = R_{\alpha\beta\mu}^{\mu} = \partial_{\beta}\Gamma_{\alpha\mu}^{\mu} - \partial_{\mu}\Gamma_{\alpha\beta}^{\mu} + \Gamma_{\alpha\mu}^{\delta}\Gamma_{\delta\beta}^{\mu} - \Gamma_{\alpha\beta}^{\delta}\Gamma_{\delta\mu}^{\mu} \quad (3)$$

- metric tensor for the + polarization

$$[g_{\mu\nu}] = [\eta^{\mu\nu}] + [h^{\mu\nu}] = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 + h(t, z) & 0 & 0 \\ 0 & 0 & -1 - h(t, z) & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad (4)$$

- inverse metric tensor for the + polarization

$$[g^{\mu\nu}] = [\eta^{\mu\nu}] - [h^{\mu\nu}] = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 - h(t, z) & 0 & 0 \\ 0 & 0 & -1 + h(t, z) & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad (5)$$

Next, assuming a sinusoidal time dependence

$$h(t, z) = A \cos(\omega t - kz) = h_{xx}^{TT} = -h_{yy}^{TT} \quad (6)$$

and using the *Diagonal metric worksheet* of T. A. Moore<sup>1</sup> we carry out the calculations that show that for the + polarization

$$R_{tt} = -\frac{1}{c^2} \left( h\ddot{h} + \frac{\dot{h}^2}{2} \right); \quad R_{zz} = R_{tt}; \quad R_{xx} = R_{yy} = 0 \quad (7)$$

to second order in  $h$ .

After averaging over time, we find

$$\begin{aligned} -\left\langle h\ddot{h} + \frac{\dot{h}^2}{2} \right\rangle &= \omega^2 A^2 \left\langle \cos^2(\omega t - kz) - \frac{1}{2} \sin^2(\omega t - kz) \right\rangle \\ &= \omega^2 A^2 \langle \cos^2(\omega t - kz) - \sin^2(\omega t - kz) \rangle + \omega^2 A^2 \left\langle \frac{1}{2} \sin^2(\omega t - kz) \right\rangle \\ &= \omega^2 A^2 \langle \cos(2\omega t - 2kz) \rangle + \frac{1}{2} \langle \dot{h}^2 \rangle = \frac{1}{2} \langle \dot{h}^2 \rangle \quad (8) \end{aligned}$$

Using equations (7) and recalling the form of the metric, we can easily prove that for the gravitational wave  $R = 0$ . Therefore

$$\langle G_{tt}^{(2)} \rangle = \langle R_{tt}^{(2)} \rangle = \frac{1}{2} \langle \dot{h}^2 \rangle \quad (9)$$

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<sup>1</sup>We must be careful that in Moore's text he uses the  $(-,+,+,+)$  signature while we use the  $(+,-,-,-)$  signature, and units such that  $c = 1$ . He also uses a different convention for the Ricci tensor, so that the sign is reversed with respect to our definition.